Predictions from analytic parametrisations at t = 0

J.R. Cudell

September 4, 2002

in collaboration with : V.V. Ezhela, P. Gauron, K. Kang, Yu.V. Kuyanov, S.B. Lugovsky, E. Martynov, B. Nicolescu, E.A. Razuvaev, N.P. Tkachenko

Predictions from analytic parametrisations at t = 0

J.R. Cudell

September 4, 2002

in collaboration with : V.V. Ezhela, P. Gauron, K. Kang, Yu.V. Kuyanov, S.B. Lugovsky, E. Martynov, B. Nicolescu, E.A. Razuvaev, N.P. Tkachenko

(COMPETE collaboration)

COmputerised Models and Parameter Evaluation for Theory and Experiment COmputerised Models and Parameter Evaluation for Theory and Experiment

Collaboration to create phenomenological knowledge bases in particle physics. So far, results for forward hadronic scattering (this talk) and $\sigma^{e^+e^-}$.

COmputerised Models and Parameter Evaluation for Theory and Experiment

Collaboration to create phenomenological knowledge bases in particle physics. So far, results for forward hadronic scattering (this talk) and $\sigma^{e^+e^-}$.

Talk based on the ideas and methods of the following papers:

- Benchmarks for the forward observables at RHIC, the Tevatron-run II and the LHC, hep-ph/0206172, accepted by Phys. Rev. Letters.
- Review of particle physics, Particle Data Group (K. Hagiwara et al.), Phys. Rev. D 66, 010001 (2002).
- Hadronic scattering amplitudes: medium-energy constraints on asymptotic behaviour, Phys. Rev. D 65, 074024 (2002).

Questions

- What are the best models describing soft forward data?
- What should be measured?
- What are the best predictions?

Questions

- What are the best models describing soft forward data?
- What should be measured?
- What are the best predictions?

Outline

- + Motivation
- + Tools
- + Results on models and data
- + Predictions



Cure the high degree of arbitrariness in the phenomenology



Cure the high degree of arbitrariness in the phenomenology

- Excessive focus on pp and $\bar{p}p$ scattering;
- Excessive focus on total cross sections;
- Fundamental physical constraints mixed with ad-hoc properties;
- Dataset varies from author to author;
- Cut-off in energy $\sqrt{s_{min}}$ varies from author to author;
- No attention paid to the stability of the parameters when some data are excluded or when $\sqrt{s_{min}}$ is varied.



Cure the high degree of arbitrariness in the phenomenology

- Excessive focus on pp and $\bar{p}p$ scattering;
- Excessive focus on total cross sections;
- Fundamental physical constraints mixed with ad-hoc properties;
- Dataset varies from author to author;
- Cut-off in energy $\sqrt{s_{min}}$ varies from author to author;
- No attention paid to the stability of the parameters when some data are excluded or when $\sqrt{s_{min}}$ is varied.

\Rightarrow Provide a common test ground for a variety of models and judge them according to the same criteria.



- Theoretical (non perturbative):
 - * analyticity, crossing symmetry, unitarity, positivity;
 - ★ Regge relation between poles and resonance masses.



- Theoretical (non perturbative):
 - * analyticity, crossing symmetry, unitarity, positivity;
 - ★ Regge relation between poles and resonance masses.
- Experimental (COMPAS database):
 - \star use both σ_{tot} and ho;
 - \star all data pp, $\bar{p}p$, $\pi^{\pm}p$, $K^{\pm}p$, $\Sigma^{-}p$, γp , $\gamma\gamma.$



- Theoretical (non perturbative):
 - * analyticity, crossing symmetry, unitarity, positivity;
 - ★ Regge relation between poles and resonance masses.
- Experimental (COMPAS database):
 - * use both σ_{tot} and ρ ; * all data pp, $\bar{p}p$, $\pi^{\pm}p$, $K^{\pm}p$, $\Sigma^{-}p$, γp , $\gamma \gamma$.

• Computer technology:

- * all fits running through a common minimization procedure under Mathematica, crossed-checked by MINUIT fortran codes;
- ★ artificial intelligence criteria;
- * web predictor at http://www.ihep.su/~tka4ehko/CS/MODELS and web interface at http://sirius.ihep.su/~kuyanov/OK/eng/intro.html.

J.R. Cudell, COMPETE collaboration

Alushta, Sept.4, 2002



Theoretical tools

Analyticity:

$$\sigma_{tot}(s) = \frac{1}{s} \Im m \left[A(s,0) \right] \Leftrightarrow \rho(s) = \frac{\Re e \left[A(s,0) \right]}{\Im m [A(s,0)]},$$

but this works only if one knows either function <u>exactly</u>. Experimental and theoretical uncertainties \rightarrow infinite number of possibilities.

Theoretical tools

Analyticity:

$$\sigma_{tot}(s) = rac{1}{s} \Im m \left[A(s,0)
ight] \Leftrightarrow
ho(s) = rac{\Re e \left[A(s,0)
ight]}{\Im m [A(s,0)]},$$

but this works only if one knows either function <u>exactly</u>. Experimental and theoretical uncertainties \rightarrow infinite number of possibilities.

Unitarity

Polynomial boundedness of absorptive part in Lehmann ellipse \Rightarrow

$$\sigma_{tot}(s) \leq C \log^2 \frac{s}{s_0} \quad as \quad s \to \infty \quad (\text{Froissart-Martin})$$

 $C \simeq \frac{\pi}{m_{\pi}^2} \simeq 60 ext{ mb} ext{ (Lukaszuk-Martin)} \Rightarrow 1 ext{ barn at the Tevatron.}$

Regge trajectories

The leading meson trajectories are seen in a Chew-Frautschi plot \Rightarrow Their intercepts can be measured directly

Regge trajectories

The leading meson trajectories are seen in a Chew-Frautschi plot \Rightarrow Their intercepts can be measured directly

Intercepts ≈ 0.5 Are the trajectories degenerate? Are they linear? Alushta, Sept.4, 2002



Positivity

All total cross sections must be positive.

Zweig's rule \rightarrow the pomeron contribution must be positive.

Small violation of Zweig's rule are possible \rightarrow only the C = +1 part of cross sections must be bigger than the C = -1 part.

Experimental data

• Need for renewed updates to the database;

Experimental data

• Need for renewed updates to the database;

Problems:

- ***** Huge gap between the ISR and the Sp \bar{p} S;
- ★ Contradictory data, *e.g.* at the Tevatron;
- **\star** Poor quality of some of the ρ data.

Alushta, Sept.4, 2002



Computer tools

Classification of models

$$\sigma^{ab}_{tot}(s) = Y^{ab}(s) + H^{ab}(s)$$

Computer tools

Classification of models

$$\sigma^{ab}_{tot}(s) = Y^{ab}(s) + H^{ab}(s)$$

• contribution Y^{ab} of the highest meson trajectories (ρ , ω , a and f)

$$Y^{ab} = Y^{ab}_{+}(s)^{\alpha_{+}-1} \pm Y^{ab}_{-}(s)^{\alpha_{-}-1} \to RR$$

Computer tools

Classification of models

$$\sigma^{ab}_{tot}(s) = Y^{ab}(s) + H^{ab}(s)$$

• contribution Y^{ab} of the highest meson trajectories (ρ , ω , a and f)

$$Y^{ab} = Y^{ab}_{+}(s)^{\alpha_{+}-1} \pm Y^{ab}_{-}(s)^{\alpha_{-}-1} \to RR$$

• rising C = +1 term H^{ab} from the pomeron contribution

$$H^{ab} = P^{ab} + X^{ab}(s)^{\alpha_{\mathcal{P}}-1} \rightarrow PE$$

$$H^{ab} = P^{ab} + B^{ab} \ln(s/s_0) \rightarrow PL$$

$$H^{ab} = P^{ab} + B^{ab} \ln^2(s/s_0) \rightarrow PL2$$

Possible constraints on the parameters:

- * degeneracy of the reggeon trajectories $(\alpha_+ = \alpha_-) \rightarrow d$
- \star universality of rising terms (B^{ab} independent of the hadrons) $\rightarrow {}_{u}$
- * factorization for $\gamma\gamma$ and γp $(H_{\gamma\gamma} = \delta H_{\gamma p} = \delta^2 H_{pp}) \rightarrow {}_{nf}$ if absent
- \star quark counting rules (Σp from pp, Kp and πp) \rightarrow $_{qc}$
- \star Johnson-Treiman-Freund relation for the cross section differencesightarrow $_c$

Possible constraints on the parameters:

- * degeneracy of the reggeon trajectories $(\alpha_+ = \alpha_-) \rightarrow d$
- \star universality of rising terms (B^{ab} independent of the hadrons)ightarrow
- * factorization for $\gamma\gamma$ and γp $(H_{\gamma\gamma} = \delta H_{\gamma p} = \delta^2 H_{pp}) \rightarrow {}_{nf}$ if absent
- \star quark counting rules (Σp from pp, Kp and πp) \rightarrow $_{qc}$
- \star Johnson-Treiman-Freund relation for the cross section differences ightarrow c

256 possibilities among which:

- Donnachie-Landshoff = $(RR)_d E$
- Cudell-Kang-Kim = RRE
- Gauron-Nicolescu = $(RR)_d PL2_u$
- Desgrolard-Giffon-Lengyel-Martynov-Predazzi= *RRPL*
- Bourrely-Soffer-Wu or other eikonals = RRL2 asymptotically

Dataset

Reaction	Number of
	data points
	for $\sqrt{s} \geq 5~{ m GeV}$
σ_{pp}	112
$\sigma_{\overline{p}p}$	59
σ_{π^+p}	50
σ_{π^-p}	106
σ_{K^+p}	40
σ_{K^-p}	63
σ_{Σ^-p}	9
$\sigma_{\gamma p}$	38
$\sigma_{\gamma\gamma}$	30
$ ho_{pp}$	74
$ ho_{\overline{p}p}$	11
$ ho_{\pi^+p}$	8
$ ho_{\pi^-p}$	30
ρ_{K^+p}	10
$ ho_{K^-p}$	8

Dataset

Reaction	Number of
	data points
	for $\sqrt{s} \geq 5~{ m GeV}$
σ_{pp}	112
$\sigma_{\overline{p}p}$	59
σ_{π^+p}	50
σ_{π^-p}	106
σ_{K^+p}	40
σ_{K^-p}	63
σ_{Σ^-p}	9
$\sigma_{\gamma p}$	38
$\sigma_{\gamma\gamma}$	30
$ ho_{pp}$	74
$ ho_{\overline{p}p}$	11
$ ho_{\pi^+p}$	8
$ ho_{\pi^-p}$	30
$ ho_{K^+p}$	10
ρ_{K^-p}	8

$\sqrt{s_{min}}$	Total number
	of data points
3 GeV	904
4 GeV	742
5 GeV	648
6 GeV	569
7 GeV	498
8 GeV	453
9 GeV	397
10 GeV	329

Criteria for A.I. decisions: ACCURRSS scheme

 $sets = \{observable \ (\sigma \ or \ \rho), beam, target \}$

A pplicability: range in energy over which a model M is valid (global fit with CL > 50%).

$$A^{M} = \frac{1}{N_{sets}} \sum_{j} w_{j}^{M} \log\left(\frac{E_{j}^{M,high}}{E_{j}^{M,low}}\right) \text{with } w_{j}^{M} = \min\left(1, \frac{1}{\chi^{2}/nop}\right);$$

Criteria for A.I. decisions: ACCURRSS scheme

 $sets = \{observable \ (\sigma \ or \ \rho), beam, target \}$

A pplicability: range in energy over which a model M is valid (global fit with CL > 50%).

$$A^{M} = \frac{1}{N_{sets}} \sum_{j} w_{j}^{M} \log\left(\frac{E_{j}^{M,high}}{E_{j}^{M,low}}\right) \text{with } w_{j}^{M} = \min\left(1, \frac{1}{\chi^{2}/nop}\right);$$

C onfidence-1: within the area of applicability of M: $C_1^M = CL(\%)$ C onfidence-2: within the considered range of energy: $C_2^M = CL(\%)$

Criteria for A.I. decisions: ACCURRSS scheme

 $sets = \{observable \ (\sigma \ or \ \rho), beam, target \}$

A pplicability: range in energy over which a model M is valid (global fit with CL > 50%).

$$A^{M} = \frac{1}{N_{sets}} \sum_{j} w_{j}^{M} \log\left(\frac{E_{j}^{M,high}}{E_{j}^{M,low}}\right) \text{with } w_{j}^{M} = \min\left(1, \frac{1}{\chi^{2}/nop}\right);$$

C onfidence-1: within the area of applicability of M: $C_1^M = CL(\%)$ C onfidence-2: within the considered range of energy: $C_2^M = CL(\%)$ U niformity: variation of the χ^2/nop from set to set:

$$U^{M} = \left\{ \frac{1}{N_{sets}} \sum_{j} \frac{1}{4w_{j}^{M}} \left[\frac{\chi^{2}}{N_{nop}} - \frac{\chi^{2}(j)}{N_{nop}^{j}} \right]^{2} \right\}^{-1}$$

R igidity: number of parameters compared to the number of data points where the model is applicable:

$$R_1^M = \frac{N_{nop}^M}{1 + N_{par}^M}$$

J.R. Cudell, COMPETE collaboration

R igidity: number of parameters compared to the number of data points where the model is applicable:

$$R_1^M = \frac{N_{nop}^M}{1 + N_{par}^M}$$

R eliability: quality of the error matrix:

$$R_2^M = \frac{2}{N_{par}(N_{par} - 1)} \cdot \sum_{i>j=1}^N \Theta(90.0 - C_{ij})$$

where C_{ij} is the correlation matrix element in % calculated in the fit at the low edge of the applicability area.

J.R. Cudell, COMPETE collaboration

R igidity: number of parameters compared to the number of data points where the model is applicable:

$$R_1^M = \frac{N_{nop}^M}{1 + N_{par}^M}$$

R eliability: quality of the error matrix:

$$R_2^M = \frac{2}{N_{par}(N_{par} - 1)} \cdot \sum_{i>j=1}^N \Theta(90.0 - C_{ij})$$

where C_{ij} is the correlation matrix element in % calculated in the fit at the low edge of the applicability area.

S tability-1: stability of the parameter values P_i when one varies the minimum energy of the fit.

$$S_{1}^{M} = \frac{N_{steps}N_{par}^{M}}{\sum_{steps,ij} (P_{i} - P_{i}^{step})(W^{t} + W^{step})_{ij}^{-1}(P_{j} - P_{j}^{step})}$$

where step = 1 GeV shift of $\sqrt{s_{min}}$ and W^t and W^{step} are the error matrices.

S tability-2: stability of the parameter values P_i when one removes the ρ data $(o = 1 \text{ with } \rho, o = 0 \text{ without}).$

$$S_2^M = \frac{2N_{par}^M}{\sum_{o,ij} (P_i - P_i^o)(W^t + W^o)_{ij}^{-1}(P_j - P_j^o)}$$
S tability-2: stability of the parameter values P_i when one removes the ρ data $(o = 1 \text{ with } \rho, o = 0 \text{ without}).$

$$S_2^M = \frac{2N_{par}^M}{\sum_{o,ij} (P_i - P_i^o)(W^t + W^o)_{ij}^{-1}(P_j - P_j^o)}$$

 $Rank \rightarrow$ number of points attributed to one model when comparing its indicators to those of the other models. Higher rank=better model.



Models

- * excluded models;
- ★ best models.



Models

- excluded models;
- ★ best models.

• Data

- ★ quality of the parts of the data sample;
- ★ the Tevatron data;
- \star the cosmic ray data.

Excluded models

 χ^2/dof , ho data excluded.

	$\sqrt{s_{min}}$ in GeV						
Model	3	4	5	6	7		
RRE	1.38	1.15	0.91	0.87	0.89		
RRPL	1.33	0.98	0.85	0.83	0.87		
RRL2	1.33	1.05	0.88	0.85	0.91		
$\mathrm{RRPL2}_{u}$	1.26	0.97	0.81	0.79	0.82		
$(\mathrm{RR})^d \mathrm{P} \mathrm{L2}_u$	1.27	0.99	0.82	0.80	0.83		

χ^2/dof , ho data included.

	$\sqrt{s_{min}}$ in GeV						
Model	4	5	6	8	10		
RRE	1.38	1.12	1.10	1.05	1.02		
RRPL	1.11	0.98	0.98	0.94	0.91		
RRL2	1.34	1.11	1.10	1.06	1.00		
$RRPL2_u$	1.14	0.97	0.97	0.92	0.92		
$(\mathrm{RR})^d \mathrm{P} \mathrm{L2}_u$	1.26	0.99	0.99	0.93	0.93		

χ^2/dof , ρ data included.

	$\sqrt{s_{min}}$ in GeV						
Model	4	5	6	8	10		
RRE	1.38	1.12	1.10	1.05	1.02		
RRPL	1.11	0.98	0.98	0.94	0.91		
RRL2	1.34	1.11	1.10	1.06	1.00		
$\mathrm{RRPL2}_{u}$	1.14	0.97	0.97	0.92	0.92		
$(\mathrm{RR})^d \mathrm{P} \mathrm{L2}_u$	1.26	0.99	0.99	0.93	0.93		

- All models based on one simple-pole are excluded by the ρ data if $\sqrt{s_{min}} \leq 10 \ GeV$. (21 models survive out of 256)

χ^2/dof , ρ data included.

	$\sqrt{s_{min}}$ in GeV						
Model	4	5	6	8	10		
RRE	1.38	1.12	1.10	1.05	1.02		
RRPL	1.11	0.98	0.98	0.94	0.91		
RRL2	1.34	1.11	1.10	1.06	1.00		
$\mathrm{RRPL2}_{u}$	1.14	0.97	0.97	0.92	0.92		
$(\mathrm{RR})^d \mathrm{P} \mathrm{L2}_u$	1.26	0.99	0.99	0.93	0.93		

- All models based on one simple-pole are excluded by the ρ data $if \sqrt{s_{min}} \leq 10 \ GeV.$ (21 models survive out of 256)
- + For $\sqrt{s_{min}} = 5$ GeV, 4 models survive: RRPL2_u, RRP_{nf}L2_u, (RR)_dPL2_u and RRPL.

Other excluded models

- String picture: J. A. Feigenbaum, P. G. Freund and M. Pigli, Phys. Rev. D 56 (1997) 2596 [hep-ph/9703296].
- Two-component pomeron: H. J. Lipkin, Phys. Rev. D 11 (1975) 1827;
 H. J. Lipkin, [hep-ph/9911259].

 χ^2/dof , ho data excluded

	ν	$\overline{/s_{min}}$	in GeV		
Model	3	5	7	9	
FFP-97	101	3.28	2.3	2.34	
Lipkin TCP	4.63	2.54	2.86	3.48	

Best model(s)

without ρ , $\sqrt{s_{min}} = 5$ GeV:

	A	C_1	C_2	U	R_1	R_2	S_1
RRE	2.6	92	81	51	25	0.88	0.18
RRPL	2.0	54	100	33	19	0.67	0.22
RRL2	2.6	98	87	85	27	0.90	0.20
$\mathbb{RRPL2}_{u}$	2.5	68	100	34	26	0.91	0.01
$(\mathrm{RR})^d \mathrm{PL2}_u$	2.5	55	100	44	28	0.88	0.11

Best model(s)

without ρ , $\sqrt{s_{min}} = 5$ GeV:

	A	C_1	C_2	U	R_1	R_2	S_1
RRE	2.6	92	81	51	25	0.88	0.18
RRPL	2.0	54	100	33	19	0.67	0.22
RRL2	2.6	98	87	85	27	0.90	0.20
$\mathbb{RRPL2}_{u}$	2.5	68	100	34	26	0.91	0.01
$(\mathrm{RR})^d \mathrm{PL2}_u$	2.5	55	100	44	28	0.88	0.11

with
$$\rho$$
, $\sqrt{s_{min}} = 9$ GeV:

	A	C_1	C_2	U	R_1	R_2	S_1	S_2
RRPL	2.3	67	82	26	29	0.75	0.21	1.14
RRL2	1.7	63	63	11	21	0.90	1.4	2.7
$\mathbb{RRPL2}_{u}$	2.2	68	85	23	30	0.90	0.22	0.10
$(\mathrm{RR})^d \mathrm{PL2}_u$	2.0	50	83	16	32	0.88	0.30	0.67

 \Rightarrow "league competition" between models, with equal weight to all indicators.

and the winners are:

(for $\sqrt{s} \ge 10$ GeV, and including ρ data.)

and the winners are:

Model Code	A	C_1	C_2	U	R_1	R_2	S_1	S_2	Rank
$\mathrm{RRPL2}_{u}$	42	26	42	42	34	28	12	4	230
$\mathrm{RRP}_{nf}\mathrm{L2}_{u}$	44	36	44	40	15	31	10	2	222
$\operatorname{RRL}_{nf}^*$	30	42	26	24	34	18	18	30	222

(for $\sqrt{s} \geq 10$ GeV, and including ρ data.)

and the winners are:

(for $\sqrt{s} \geq 10$ GeV, and including ρ data.)

Model Code	A	C_1	C_2	U	R_1	R_2	S_1	S_2	Rank
$\mathrm{RRPL2}_{u}$	42	26	42	42	34	28	12	4	230
$\mathrm{RRP}_{nf}\mathrm{L2}_{u}$	44	36	44	40	15	31	10	2	222
RRL_{nf}^*	30	42	26	24	34	18	18	30	222
$(\mathrm{RR}_c)^d \mathrm{PL2}_u$	34	20	36	20	28	24	28	14	204
$(\mathrm{RR})^d \mathrm{PL2}_u$	40	8	40	22	34	22	16	12	194
$\mathbf{R}^{qc}\mathbf{R}_{c} \mathbf{L}^{qc}$	14	32	18	10	42	6	24	38	184
$(\mathrm{RR}_c)^d \mathrm{P}^{qc} \mathrm{L2}_u$	20	16	10	36	19	36	22	22	181
$(\mathrm{RR})^d \mathrm{P}^{qc} \mathrm{L2}_u$	18	14	8	38	8	38	30	26	180
$\mathrm{RR}_{c} \mathrm{L2}^{qc}$	6	30	6	4	6	44	44	40	180
$(\mathrm{RR})^d \mathrm{PL2}^*$	38	2	28	32	25	31	14	8	178
$(\mathrm{RR})^d \operatorname{PL2}_u$	36	0	34	18	30	26	20	10	174
RRPL^*	2	34	32	44	15	16	6	24	173
$\operatorname{RR}_{c} \operatorname{L}^{qc}$	24	38	24	8	10	4	32	32	172
$\mathrm{RRL2}^{qc}$	10	28	4	2	2	42	40	42	170
$\mathbf{R}^{qc}\mathbf{R}_{c}$ $\mathbf{L}2^{qc}$	12	18	0	6	22	40	38	34	170
RRL^{qc}	28	6	20	30	44	12	4	18	162
RRPE_u	22	44	12	16	4	20	34	6	158
$\mathrm{R}^{qc}\mathrm{RL}^{qc}$	16	24	14	12	19	14	36	20	155
RRL2	8	22	2	0	0	34	42	44	152
RR _c PL	4	12	38	14	12	0	26	36	142
RRL	26	10	16	26	39	8	8	0	133

Quality of the dataset: $\chi^2/point$

Reaction/Model	RRPL2u	RRPL	RRE
σ_{pp}	0.872	0.866	0.889
$\sigma_{\overline{p}p}$	1.20	1.01	1.12
σ_{π^+p}	0.785	0.779	1.43
σ_{π^-p}	0.888	0.895	0.883
σ_{K^+p}	0.706	0.723	1.01
σ_{K^-p}	0.614	0.619	0.719
$\sigma_{\Sigma^- p}$	0.376	0.376	0.385
$\sigma_{\gamma p}$	0.602	0.752	0.586
$\sigma_{\gamma\gamma}$	0.517	0.947	0.552
$ ho_{pp}$	1.74	1.57	1.76
$ ho_{\overline{p}p}$	0.548	0.468	0.599
$ ho_{\pi^+p}$	1.45	1.59	2.71
ρ_{π^-p}	1.16	1.268	2.11
ρ_{K^+p}	1.16	1.11	0.833
ρ_{K^-p}	0.966	1.24	1.77

No model can fit the real part of the pp and πp amplitudes (see next talk by Selyugin). These data are those that exclude simple poles.

The Tevatron data

 χ^2/dof using the database of the 2002 Review of Particle Physics +new ZEUS data + best model RRPL2 $_u$.

The Tevatron data

 χ^2/dof using the database of the 2002 Review of Particle Physics +new ZEUS data + best model RRPL2_u.

Data with	all	E710/E811	CDF					
a change in χ^2		only	only					
total	0.966	0.964	0.951					
total cross sections								
$\overline{p}p$	1.15	1.12	1.05					
K^-p	0.62	0.62	0.61					
$\gamma\gamma$	0.64	0.64	0.63					
elastic forward Re/Im								
$\overline{p}p$	0.52	0.52	0.53					
pp	1.83	1.83	1.80					
$\pi^- p$	1.10	1.09	1.14					
$\pi^+ p$	1.50	1.52	1.46					
K^-p	0.99	1.01	0.96					
K^+p	1.07	1.10	0.98					
values of the parameter B								
	0.307(10)	0.301(10)	0.327(10)					

The Tevatron data

 χ^2/dof using the database of the 2002 Review of Particle Physics +new ZEUS data + best model RRPL2_u.

Data with	all	E710/E811	CDF		
a change in χ^2		only	only		
total	0.966	0.964	0.951		
	total cross	sections			
$\overline{p}p$	1.15	1.12	1.05		
K^-p	0.62	0.62	0.61		
$\gamma\gamma$	0.64	0.64	0.63		
	elastic forward Re/Im				
$\overline{p}p$	0.52	0.52	0.53		
pp	1.83	1.83	1.80		
$\pi^- p$	1.10	1.09	1.14		
$\pi^+ p$	1.50	1.52	1.46		
K^-p	0.99	1.01	0.96		
K^+p	1.07	1.10	0.98		
V	values of the parameter B				
	0.307(10)	0.301(10)	0.327(10)		

 $\Rightarrow Preference for the CDF data. Similar conclusion in the case of simple poles.$

Cosmic ray data

Data samples:

- original experimental (R. M. Baltrusaitis *et al.*, Phys. Rev. Lett. **52** (1984) 1380; M. Honda *et al.*, Phys. Rev. Lett. **70** (1993) 525.);
- corrected by Nikolaev et al. (B. Z. Kopeliovich, N. N. Nikolaev and I. K. Potashnikova, Phys. Rev. D 39 (1989) 769; N. N. Nikolaev, Phys. Rev. D 48 (1993) 1904 [hep-ph/9304283].);
- corrected by Block et al. and Durand (L. Durand and H. Pi, Phys. Rev. D 38 (1988) 78; M. M. Block, F. Halzen and T. Stanev, Phys. Rev. D 62 (2000) 077501 [hep-ph/0004232]).

Cosmic ray data

Data samples:

- original experimental (R. M. Baltrusaitis *et al.*, Phys. Rev. Lett. **52** (1984) 1380; M. Honda *et al.*, Phys. Rev. Lett. **70** (1993) 525.);
- corrected by Nikolaev et al. (B. Z. Kopeliovich, N. N. Nikolaev and I. K. Potashnikova, Phys. Rev. D 39 (1989) 769; N. N. Nikolaev, Phys. Rev. D 48 (1993) 1904 [hep-ph/9304283].);
- corrected by Block et al. and Durand (L. Durand and H. Pi, Phys. Rev. D 38 (1988) 78; M. M. Block, F. Halzen and T. Stanev, Phys. Rev. D 62 (2000) 077501 [hep-ph/0004232]).

			+30 mb		-20 mb	
	Exp	eriment	Nikolaev et al.		Block et al.	
Model	χ^2	χ^2/N_{dp}	χ^2	χ^2/N_{dp}	χ^2	χ^2/N_{dp}
$\mathrm{RRPL2}_{u}$	1.62	0.23	14.31	2.04	3.30	0.47
RRPL	2.93	0.42	25.56	3.64	2.34	0.33
RRE	1.73	0.25	14.60	2.1	3.45	0.49

Cosmic ray data

Data samples:

- original experimental (R. M. Baltrusaitis *et al.*, Phys. Rev. Lett. **52** (1984) 1380; M. Honda *et al.*, Phys. Rev. Lett. **70** (1993) 525.);
- corrected by Nikolaev et al. (B. Z. Kopeliovich, N. N. Nikolaev and I. K. Potashnikova, Phys. Rev. D 39 (1989) 769; N. N. Nikolaev, Phys. Rev. D 48 (1993) 1904 [hep-ph/9304283].);
- corrected by Block et al. and Durand (L. Durand and H. Pi, Phys. Rev. D 38 (1988) 78; M. M. Block,
 F. Halzen and T. Stanev, Phys. Rev. D 62 (2000) 077501 [hep-ph/0004232]).

			+30 mb		-20 mb	
	Exp	eriment	riment Nikolaev et al.		Block et al.	
Model	χ^2	χ^2/N_{dp}	χ^2	χ^2/N_{dp}	χ^2	χ^2/N_{dp}
$\mathrm{RRPL2}_{u}$	1.62	0.23	14.31	2.04	3.30	0.47
RRPL	2.93	0.42	25.56	3.64	2.34	0.33
RRE	1.73	0.25	14.60	2.1	3.45	0.49

 \Rightarrow original experimental analysis favoured



- RHIC
- Tevatron run II
- LHC
- cosmic rays

RHIC, Tevatron Run II and the LHC

	$\sqrt{s}~({\sf GeV})$	$\sigma~({\sf mb})$	ρ
	100	$46.37 \pm 0.06 \begin{array}{c} +0.17 \\ -0.09 \end{array}$	$\begin{array}{r} 0.1058 \pm 0.0012 & +0.0040 \\ -0.0021 \end{array}$
	200	$51.76 \pm 0.12 \begin{array}{c} +0.39 \\ -0.21 \end{array}$	$\begin{array}{r} 0.1275 \pm 0.0015 & +0.0051 \\ -0.0026 \end{array}$
	300	$55.50 \pm 0.17 \begin{array}{c} +0.57 \\ -0.30 \end{array}$	$0.1352 \pm 0.0016 \begin{array}{c} +0.0055 \\ -0.0028 \end{array}$
$RRPL2_u \ value$	400	$58.41 \pm 0.21 \begin{array}{c} +0.71 \\ -0.36 \end{array}$	$0.1391 \pm 0.0017 \begin{array}{c} +0.0056 \\ -0.0030 \end{array}$
$\pm statistical$	500	$60.82 \pm 0.25 \begin{array}{c} +0.82 \\ -0.45 \end{array}$	$0.1413 \pm 0.0017 \begin{array}{c} +0.0057 \\ -0.0030 \end{array}$
$\pm Tevatron$ disagreement	600	$62.87 \pm 0.28 \begin{array}{c} +0.94 \\ -0.48 \end{array}$	$0.1416 \pm 0.0018 egin{array}{c} +0.0058 \ -0.0031 \end{array}$
ateagreentent	1960	$78.27 \pm 0.55 \begin{array}{c} +1.85 \\ -0.96 \end{array}$	$0.1450 \pm 0.0018 \begin{array}{c} +0.0057 \\ -0.0030 \end{array}$
	10000	$105.1 \pm 1.1 \begin{array}{c} +3.6 \\ -1.9 \end{array}$	$0.1382 \pm 0.0016 \begin{array}{c} +0.0047 \\ -0.0027 \end{array}$
	12000	$108.5 \pm 1.2 \begin{array}{c} +3.8 \\ -2.0 \end{array}$	$0.1371 \pm 0.0015 \begin{array}{c} +0.0046 \\ -0.0026 \end{array}$
	14000	$\begin{array}{c} 2.0 \\ +4.1 \\ -2.1 \end{array}$	$0.1361 \pm 0.0015 \begin{array}{c} +0.0026 \\ +0.0058 \\ -0.0025 \end{array}$

RHIC, Tevatron Run II and the LHC

	$\sqrt{s}~({ m GeV}) ~\sigma~({ m mb})$		ho	
	100	$46.369 \pm 0.068 \begin{array}{c} +0.301 \\ -0.047 \end{array}$	$0.1047 \pm 0.0013 \begin{array}{c} +0.0034 \\ -0.0007 \end{array}$	
	200	$51.70 \pm 0.13 \begin{array}{c} +0.48 \\ -0.08 \end{array}$	$\begin{array}{c} 0.1260 \pm 0.0017 \\ -0.0006 \end{array} + \begin{array}{c} +0.0008 \\ -0.0006 \end{array}$	
	300	$55.39 \pm 0.18 \begin{array}{c} +0.49 \\ -0.08 \end{array}$	$0.1335 \pm 0.0019 \begin{array}{c} +0.0013 \\ -0.0039 \end{array}$	
	400	$58.25 \pm 0.22 \begin{array}{c} +0.43 \\ -0.06 \end{array}$	$0.1373 \pm 0.0021 \begin{array}{c} +0.0016 \\ -0.0065 \end{array}$	
Theoretical error	500	$60.62 \pm 0.26 \ \begin{array}{c} +0.34 \\ -0.04 \end{array}$	$0.1395 \pm 0.0022 egin{array}{c} +0.0019 \ -0.0086 \end{array}$	
	600	$62.64 \pm 0.30 \begin{array}{c} +0.24 \\ -0.01 \end{array}$	$0.1409 \pm 0.0023 \begin{array}{c} +0.0021 \\ -0.0102 \end{array}$	
	1960	$77.78 \pm 0.63 \begin{array}{c} +0.48 \\ -1.39 \end{array}$	$0.1435 \pm 0.0027 \begin{array}{c} +0.0028 \\ -0.0202 \end{array}$	
	10000	$104.1 \pm 1.4 \begin{array}{c} +1.3 \\ -7.0 \end{array}$	$0.1368 \pm 0.0028 \begin{array}{c} +0.0030 \\ -0.0298 \end{array}$	
	12000	$107.5 \pm 1.5 \begin{array}{c} +1.4 \\ 7.9 \end{array}$	$0.1358 \pm 0.0028 \begin{array}{c} +0.0030 \\ 0.0306 \end{array}$	
	14000	$110.4 \pm 1.6 \begin{array}{c} +1.5 \\ -8.7 \end{array}$	$0.1348 \pm 0.0028 \begin{array}{c} -0.0300 \\ +0.0030 \\ -0.0311 \end{array}$	



theoretical uncertainty from 21 allowed models Tevatron uncertainty

statistical uncertainty



theoretical uncertainty from 21 allowed models Tevatron uncertainty

statistical uncertainty

Cosmic rays

σ_{tot} for $\gamma p \rightarrow hadrons$, RRPL2_u

p_{lab}^{γ} (GeV)	$\sigma~({\sf mb})$		
$1.0\cdot 10^6$	0.262 ± 0.010	+0.013 -0.011	
$1.0\cdot 10^7$	0.333 ± 0.016	+0.021 -0.017	
$1.0\cdot 10^9$	0.516 ± 0.029	+0.042 -0.032	

 σ_{tot} for $\gamma\gamma \rightarrow hadrons$, RRPL2_u

$\sqrt{s}~({ m GeV})$	$\sigma~(\mu$ b)		
300	0.610 ± 0.035	+0.037	
		-0.035	
500	0.700 ± 0.047	+0.050 -0.048	
		+0.043	
1000	0.840 ± 0.067	-0.069	



Data Models













★ COMPETE project works and is almost fully implemented for forward observables;
- COMPETE project works and is almost fully implemented for forward observables;
- Using the current database, it seems impossible to decide on the singularity structure at present. However, a double-component pomeron is favoured, with a universal rising component.

- COMPETE project works and is almost fully implemented for forward observables;
- * Using the current database, it seems impossible to decide on the singularity structure at present. However, a double-component pomeron is favoured, with a universal rising component. Simple-pole models are disfavoured, among others.

- COMPETE project works and is almost fully implemented for forward observables;
- * Using the current database, it seems impossible to decide on the singularity structure at present. However, a double-component pomeron is favoured, with a universal rising component. Simple-pole models are disfavoured, among others.
- * There are problems with some sub-sets of the data: ρ , SELEX.

- COMPETE project works and is almost fully implemented for forward observables;
- * Using the current database, it seems impossible to decide on the singularity structure at present. However, a double-component pomeron is favoured, with a universal rising component. Simple-pole models are disfavoured, among others.
- **\star** There are problems with some sub-sets of the data: ρ , SELEX.
- * Predictions include an assessment of systematic errors due to experimental and theoretical disagreements.

J.R. Cudell, COMPETE collaboration

Alushta, Sept.4, 2002



Plans for the future

- Further automatisation of procedure and integration of various parts into one object of knowledge;
- Solution of the ρ problem, inclusion of the correlations between σ_{tot} and ρ ;
- Proper treatment of systematic errors;
- Link to other OKs $\leftrightarrow \rho$
 - electromagnetic form factors ↔ Coulomb interference region;
 - Regge trajectories ↔ subleading trajectories;
 - Elastic scattering.